



Roman Coronagraph Instrument
Deep Dive Technical Series

LOWFSSim

Integrated Model of LOWFS

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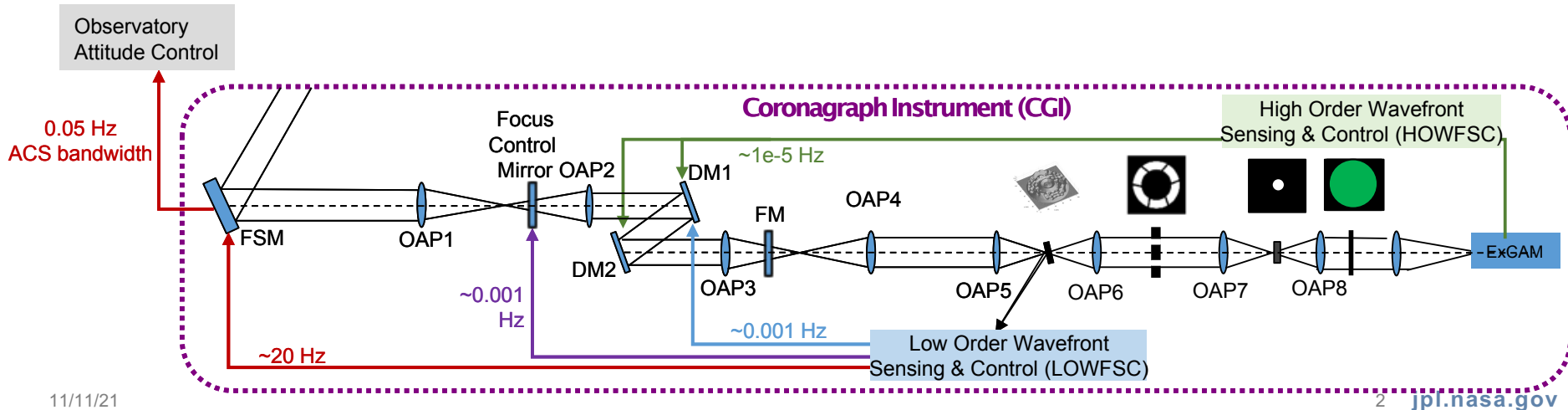
CL 21-5732

What is LOWFS?

LOWFS is the Low Order Wavefront Sensing and Control system aboard Roman-CGI.

LOWFS incorporates a **modal** Zernike Wavefront Sensor (ZWFS). It senses Z2-Z11 in Noll indexed Zernike polynomials.

It has ten control loops, spanning tip/tilt up to primary spherical aberration.



CGI Acronyms

LOWFS = Low-Order Wavefront Sensing & Control

HLC = Hybrid Lyot Coronagraph

SPC = Shaped Pupil Coronagraph

NFOV = Narrow Field-of-View

WFOV = Wide Field-of-View

LOS = Line-of-Sight

EMCCD = Electron Multiplying Charged Coupled Device

Why LOWFSSim?

What if I want to know what's happening at the full **1kHz** that hardware LOWFS runs at?

What if I want to do a study of multiple parameters and their influence on performance?

What if I want to iterate my analysis quickly; changing inputs based on what I'm seeing in real-time?

To do these things, the model has to be **fast**.

For the answers to be meaningful, the model has to be **accurate**.

What is LOWFSSim?

LOWFSSim is a model of the LOWFS system that includes:

- Diffraction
- Thin Film effects and Polarization at the occulter
- Radiometry
- Detector Effects
- Wavefront Sensing (exactly the flight algorithm, but not the flight implementation)
- Wavefront Control*
- All CGI configurations (HLC-NFOV, SPC-SPEC, SPC-WFOV)

The detector model is based on *Hirsch et al* 2013 <https://doi.org/10.1371/journal.pone.0053671>

Public repository: <https://github.com/nasa-jpl/lowfssim>

LOWFSSim is based on <https://github.com/brandondube/prysm>

Wavefront Control*

The public release of LOWFSSim includes the necessary building blocks to perform wavefront control but does not include the “prescription” for the flight controllers. Users of public LOWFSSim must design their own based on the top-level parameters:

Channels	Bandwidth (Hz)	Sample Rate (Hz)
Z2/Z3	20	1000
Z4-Z11	0.0016	0.1

Among the examples is a notebook that models closed-loop Z2, Z3, but the controller initialization is redacted and variables used are not present in the public source code.

The 0.1Hz system is derived from averages of the 1000Hz Zernike coefficient estimates; the estimator is not a multi-rate system, only the controllers are.

Wavefront Control*

A DM model is not included. The model supports injection by any phase/amplitude screen at the DM1 and DM2 planes. A user can plug in their own DM model this way, or utilize the idealized phase errors built into the model.

The omission of the actual flight controllers in the public release is not a request for the community to contribute new control designs.

Installing LOWFSSim

Download or git clone

then `pip install -e .` in the main lowfssim directory

To use a GPU, install the relevant version of cupy for your CUDA version

nasa-jpl / lowfssim Public

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main 1 branch 0 tags

Go to file Code

File/Folder	Description	Last Commit
bdube-jpl	typo in name	2235153 on Aug 31 9 commits
data	+ data files	2 months ago
docs	+ docs	2 months ago
examples	typo in name	2 months ago
lowfsc	+ actual model code	2 months ago
.gitattributes	+ data files	2 months ago
.gitignore	setup files, ignore mac DS_Store files	2 months ago
LICENSE	Update LICENSE	2 months ago
setup.cfg	update setup cfg	2 months ago
setup.py	setup files, ignore mac DS_Store files	2 months ago

About
Roman-CGI optical model of LOWFS
Apache-2.0 License

Releases
No releases published

Packages
No packages published

Languages
Python 100.0%

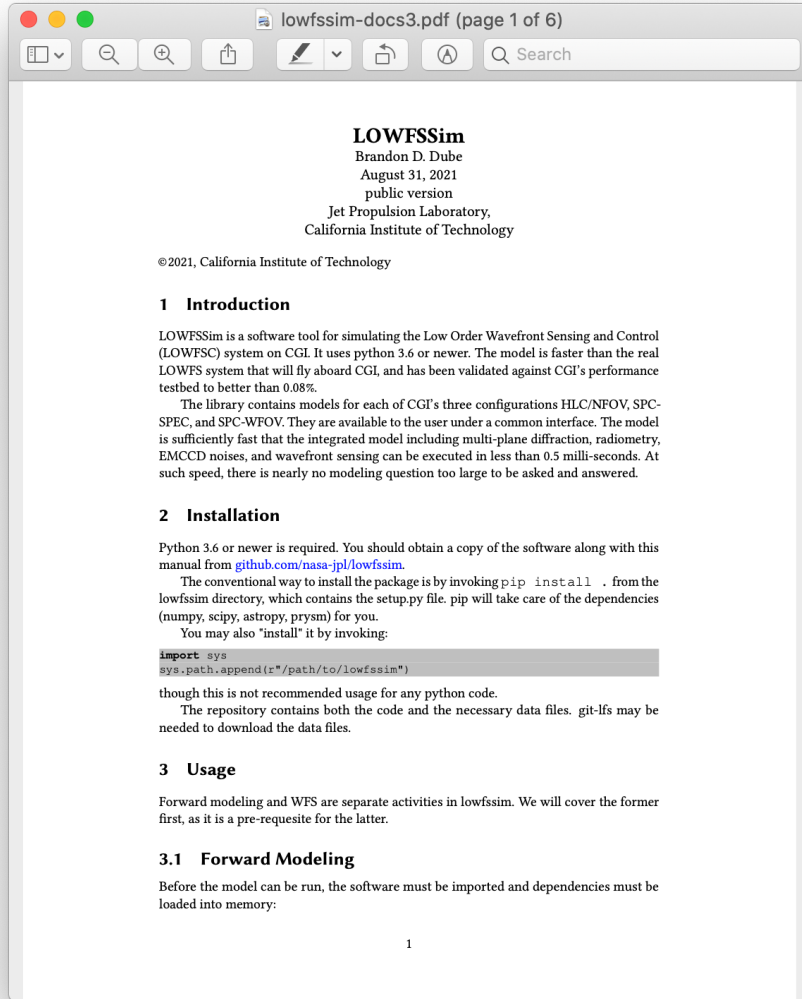
More Docs

```
# props.py
def polychromatic(wvls, weights, data, zernikes=None,
                  locam_misfocus=0, locam_shear=(0, 0), pool=None):
    """Polychromatic forward model of LOWFS.
```

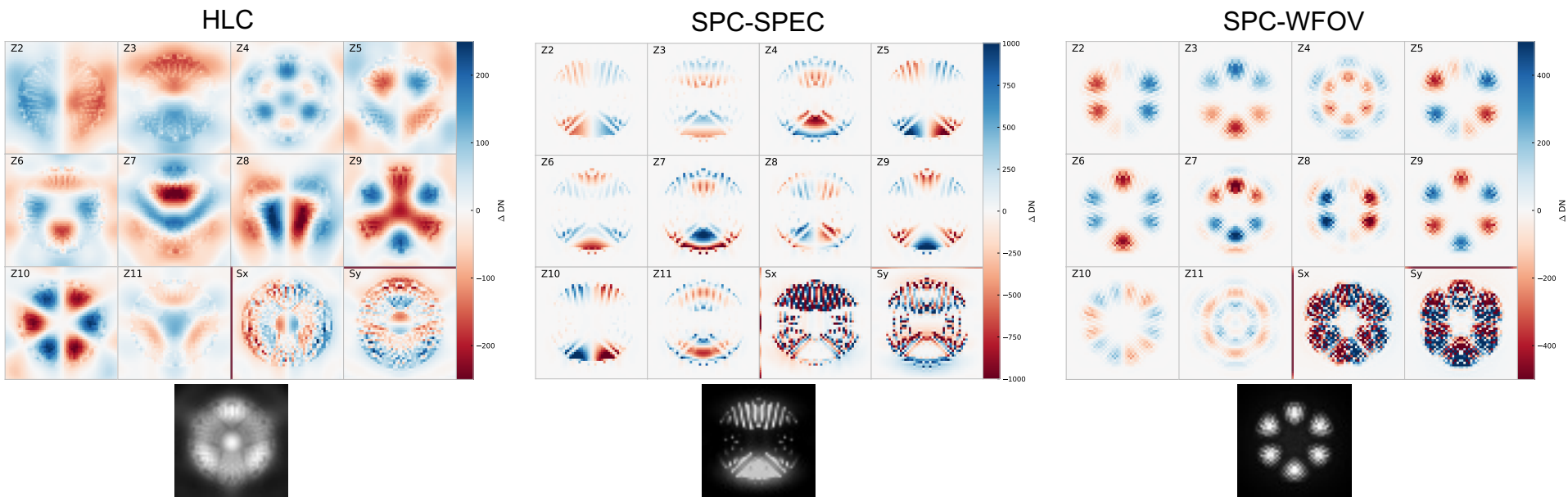
Parameters

```
-----
wvls : iterable of float
       wavelengths of light, microns
weights : iterable of float
          spectral weights to apply to the data
data : data.DesignData
       object with properties of:
       roman_pupil - amplitude map of roman pupil
       dm1_wfe - phase map in nm from DM1
       dm2_wfe - phase map in nm from DM2
       fpm - callable with a single argument of wavelength
              (in um) that produces a complex reflection map of the FPM
       pupil_mask - amplitude map of a pupil mask (i.e., for SPC)
zernikes : ndarray
           vector of Zernike coefficients, same length as dd.seed_zernikes
locam_misfocus : float, optional
                 misfocus or misconjugation of the LOWFS camera, millimeters
locam_shear : tuple of (int, int)
              shear in x and y of the image on locam, in oversampled space.
              The model has 8x oversampling at locam, so locam_shear=(1,0)
              shears by 1/8px on the output grid
pool : mapper
      a type which has a map method, multiprocessing thread and
      process pools work, as do concurrent futures equivalents
```

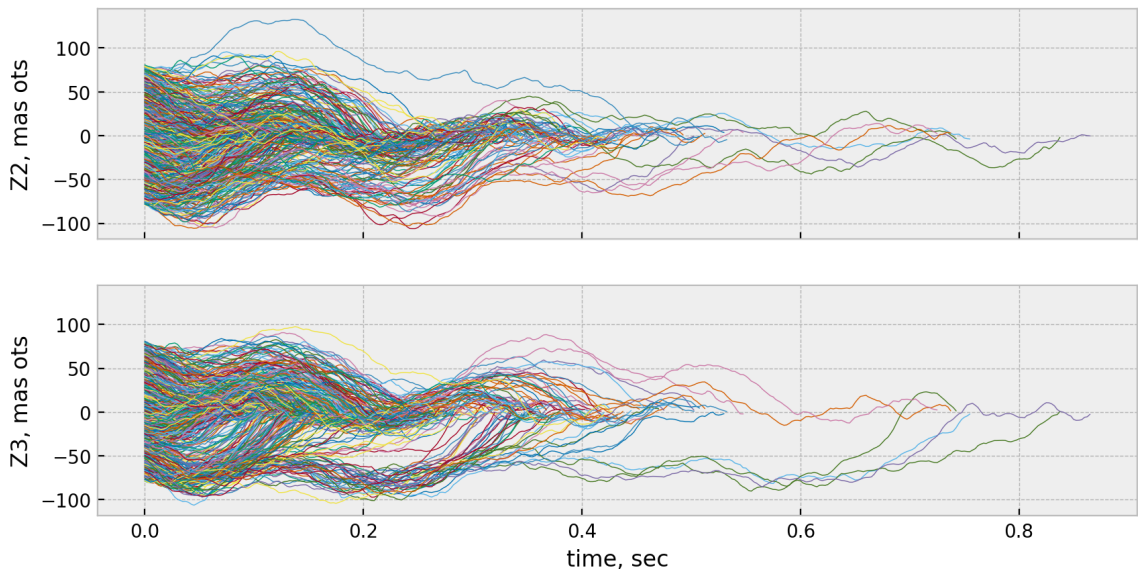
```
"""
```



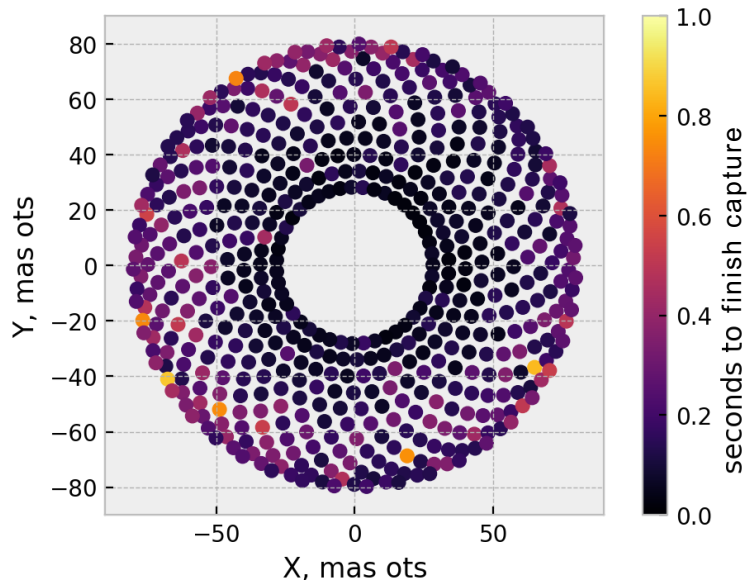
Morphological Response at LOCAM to Zernike inputs



Closed loop LOS capture and control under dynamic perturbation



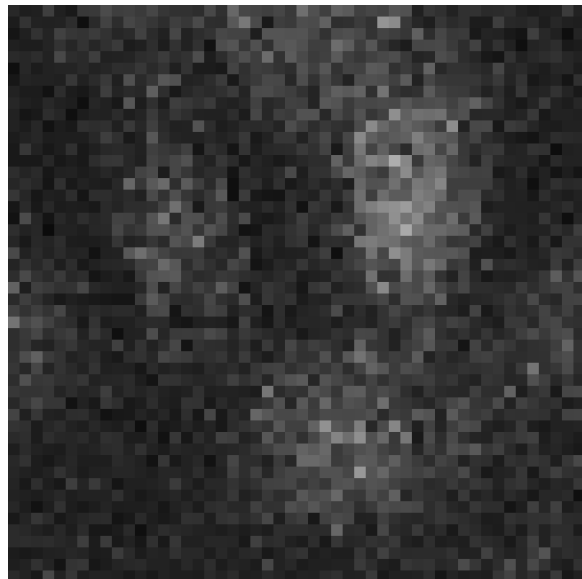
*559 trials super-imposed; millions of model evaluations
lowfssim can complete about 3 of these trials per second in
real time*



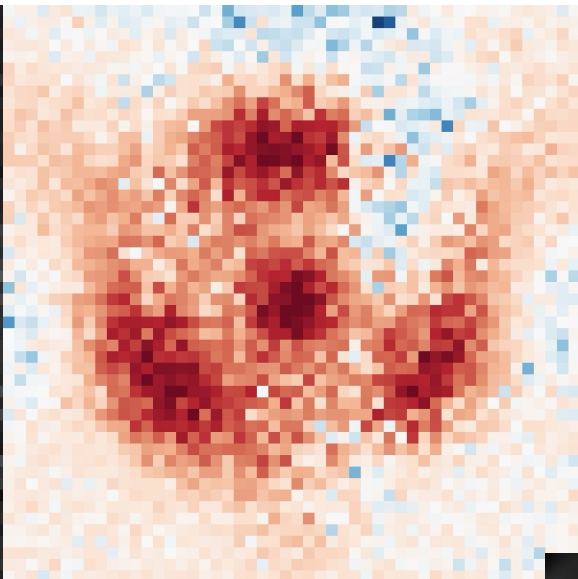
*The starting points, color
coded by how long it took for
LOS capture to conclude*

Timeseries of LOCAM frames during Target Star capture

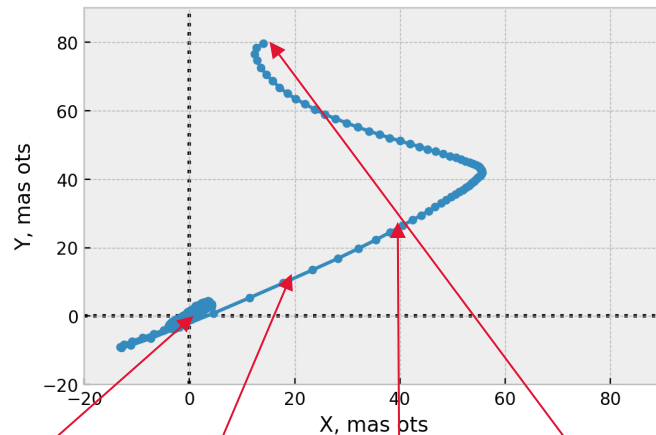
Raw Images



Differential Images



Followed LOS trajectory



The ideal differential image is pure white

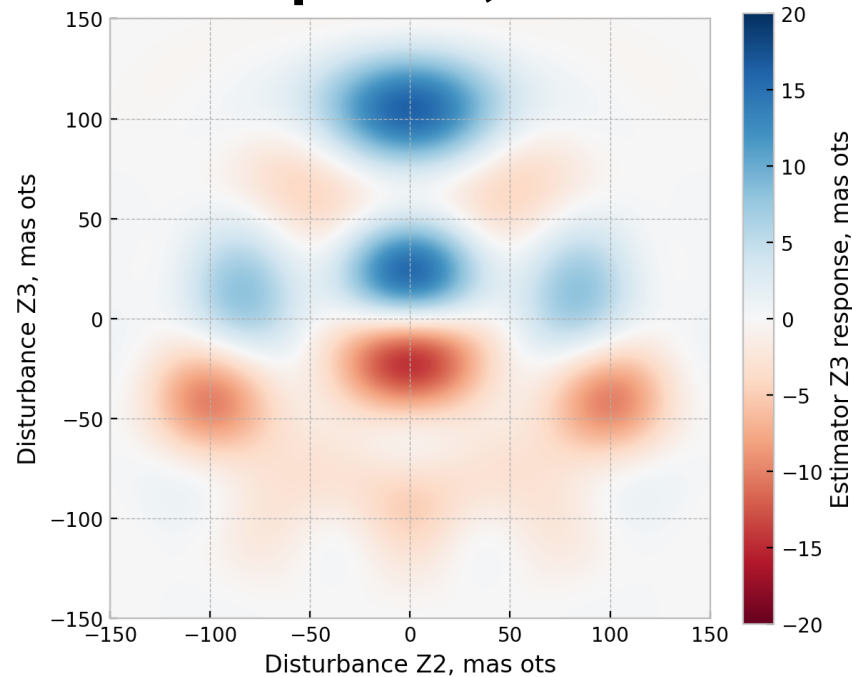
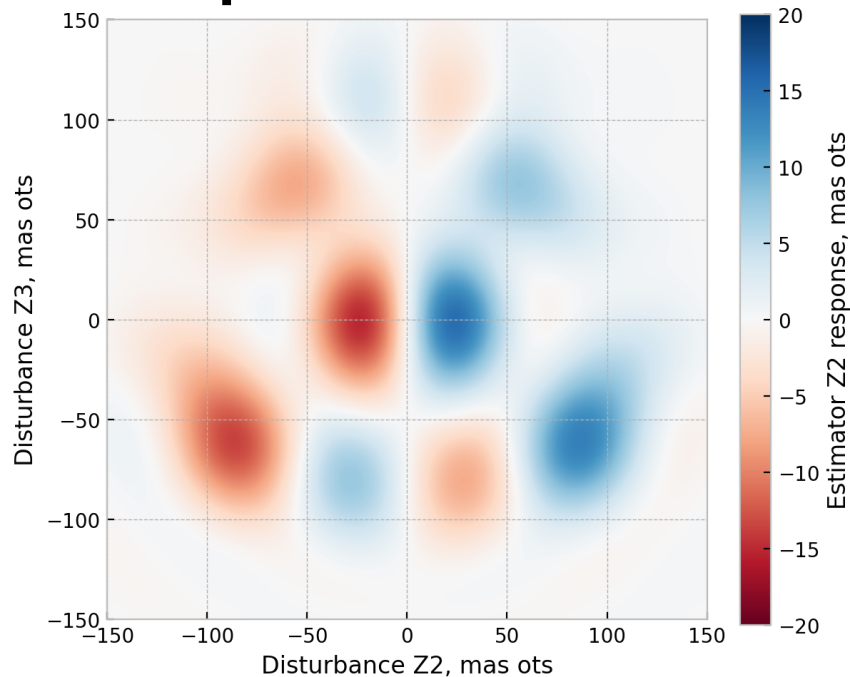
the Target Star is 47 Uma

11/11/21 *These videos advances 80x slower than real-time*



time

2D Response of LOWFS Estimator to Input Z2, Z3



These graphics show the response of the LOWFS estimator in Z2 and Z3 to a grid of real Z2, Z3 inputs. It shows the unusual response of the estimator over a large dynamic range ($\sim 3/4\lambda$ RMS).

Intermission

The proceeding slides introduce LOWFS and LOWFSSim and some interesting uses of the model.

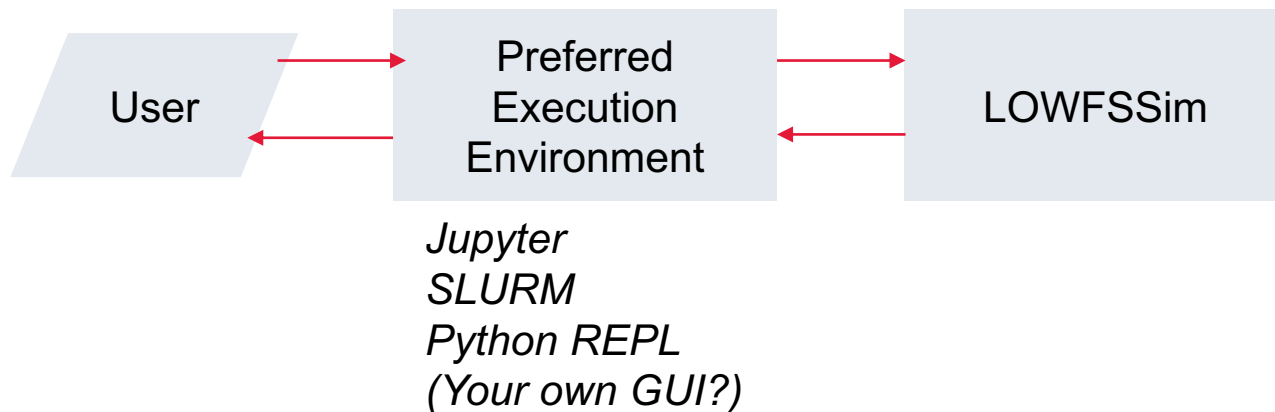
The following slides are something of a tutorial on using the model.

Any **Questions** on the content up-to-now?

Examples

1. A 2D map of the Z2, Z3 estimates over an extended dynamic range
2. Reference/Target star offsets
3. Line of Sight capture and closed-loop control of Z2/Z3

Sidebar: the User's Environment



- ✓ Laptop
- ✓ Workstation
- ✓ Server
- ✓ Supercomputer
- ✓ Interactive execution
- ✗ Define-then-run config files (but you could wrap LOWFSSim with config files)
- ✗ High level of ceremony involved in running the model

Driving LOWFSim: setup

This wall of imports reflects that part of lowfssim's speed comes from bringing everything to the top level; there is no system inside the model that forces the pieces to go together in a certain way.

These lines of code configure everything for the GPU and single precision floats; for CPU, skip them

```

from pathlib import Path

import numpy as np

from lowfsc import props, control
from lowfsc.data import DesignData
from lowfsc.spectral import StellarDatabase, LOWFS_BANDPASS, ThroughputDatabase
from lowfsc.automate import flt_chop_seq
from lowfsc.emccd import EMCCD

from matplotlib import pyplot as plt

from prysm.conf import config
from prysm.fttools import mdft

import cupy as cp
from cupyx.scipy import fft as cpfft

from prysm.mathops import np as pnp, fft as pfft

mdft.clear()
pfft._srcmodule = cpfft
pnp._srcmodule = cp
config.precision = 32

root = Path('~/.proj/wfirst/lowfs/data')
```

Driving LOWFSSim: calibration

...still not quite done loading things; this creates the camera and box that holds the DM phase screens, Roman pupil, the focal plane mask, etc. “wt” is the Zernike weights, and will be reused over and over again

This chops the estimator with 5 nm each of Z2..Z11. The chop “step” size is important to LOWFS performance and should not be changed. Non-diagonal elements of this matrix can be used to move multiple Zernike modes at once during calibration and model how these produce sensing errors.

Gains are the scalars used to calibrate each Zernike mode’s linear responsivity.

refz is where in Zernike space the estimator is calibrated to and becomes the zero point for differential estimates and control.

refst and targst are the reference and target stars.

(Vmag, spectral type, EM gain)

wref, wtarg are the spectral weights for reference and target stars. R is the LOWFS estimator.

```
dd = DesignData.hlc_design(root)
cam = EMCCD.cgi_camera()
dd.seed_zernikes(range(2,12))
wt = np.zeros(10)
```

```
refz = wt.copy()
refz[:] = 0
chops = cp.eye(10, dtype=refz.dtype)*5
gains = cp.diag(chops)
```

```
refst = (2.25, 'g0v', 20)
targst = (5, 'g0v', 150)

d = flt_chop_seq(wvl, dd, refz, chops, gains,
ref=refst, targ=targst, cam=cam)

wref, wtarg, R = d['wref'], d['wtarg'], d['R']
```

Driving LOWFSSim: images, estimates

A useful test that you have not made a mistake is to make sure the estimator recognizes its chops as themselves (5 nm, small off-diagonal terms).

The estimator returns 17 things in a vector, python indices 1:11 are Z2..Z11. The PDF docs enumerate the meaning of each index.

Immediately below for, im has units of e-/sec. After cam.expose it is an average of 10,000 images with units of DN.

If you look at the source code, props.polychromatic does very little other than the actual mechanics of diffraction. It does not draw any masks, compute any grids, etc. This is a significant component of lowfssim's speed.

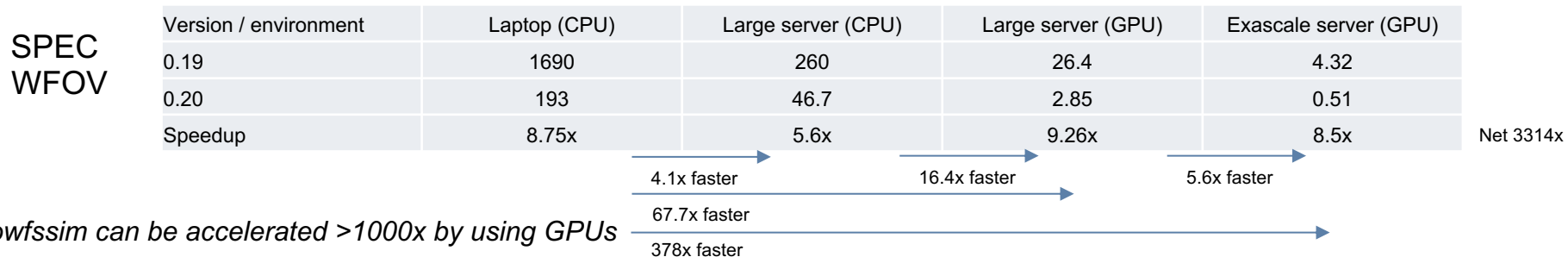
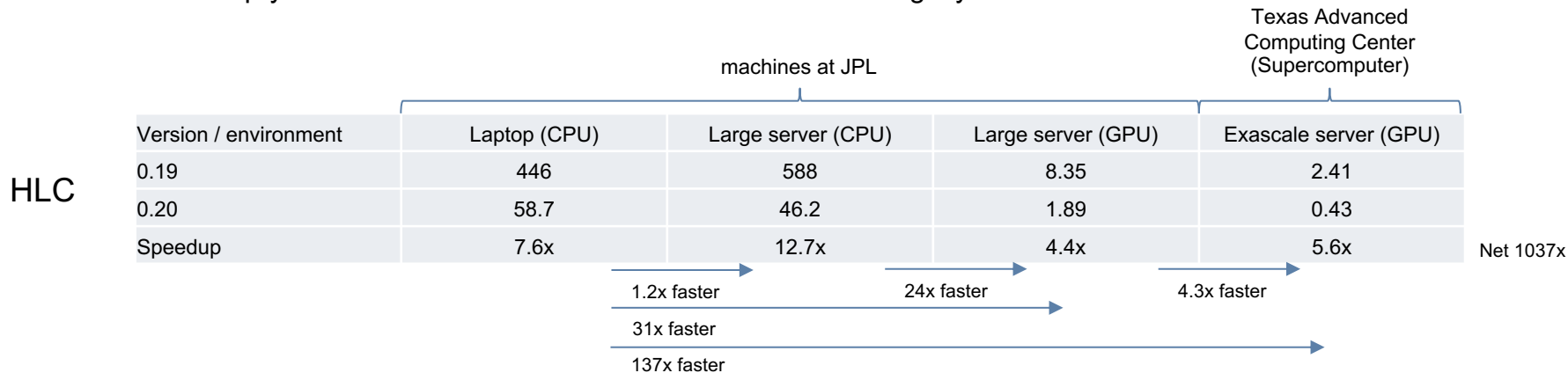
The key computation performance principle used by LOWFSSim is “load inputs once, run many times”

```
np.set_printoptions(suppress=True, precision=2, linewidth=120)
cam.em_gain = targst[2]
for i, row in enumerate(chops):
    im = props.polychromatic(wvl, wtarg, dd, zernikes=row)
    im = cam.expose(im, 10_000).mean(axis=0, dtype=im.dtype)
    e = R.estimate(im)
    resprow = e[1:11]
    print(resprow)
```

[4.97	0.01	0.08	-0.02	-0.	0.05	0.01	-0.04	0.01	0.06]
[-0.01	4.94	-0.04	-0.02	0.01	-0.02	0.02	0.	0.02	0.03]	
[0.01	0.06	4.79	-0.03	-0.01	0.04	-0.03	0.06	0.01	0.14]
[0.03	0.07	-0.05	4.97	0.03	-0.	-0.	0.04	-0.02	-0.]
[0.08	-0.07	0.05	0.03	4.95	0.	0.01	0.02	-0.04	0.05]
[0.03	-0.02	-0.01	-0.01	0.12	5.04	-0.01	0.05	-0.02	0.11]
[0.01	0.01	-0.01	-0.01	-0.12	-0.03	5.	0.09	-0.03	0.1]
[-0.01	0.	0.04	-0.01	-0.01	0.05	-0.01	5.01	0.	0.07]	
[-0.04	-0.06	0.06	0.02	-0.04	0.05	-0.02	0.	5.	-0.06]	
[-0.	0.01	0.08	-0.01	0.02	0.05	-0.05	0.05	-0.01	4.96]	

Computational Performance

On an Nvidia A100 GPU, LOWFSSim is able to model the system in closed loop at ~2.2kHz in real-time. **All times in ms.** 0.19 and 0.20 refer to prysm versions. Public lowfssim = v0.20. 0.19 is a legacy version at JPL.



lowfssim can be accelerated >1000x by using GPUs

Conclusions

LOWFSSim is a high fidelity model of LOWFS and represents the state of the art in diffraction modeling and is now available to the public to use and modify to model the next great observatories.

LOWFS itself represents the state of the art in low order wavefront sensing and control.

LOWFS will soon be in space, stabilizing CGI and enabling exoplanet science.

Questions?



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